# Physics GRE Review Fall 2004 <br> General Relativity Problems 

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## Special Relativity Problem Set

The standard MIT relativity curriculum, either from 8.033 or 8.20 , far exceeds the level of detail tested on the GRE. This problem set is intended to refresh the essential concepts you can expect to encounter on the Physics GRE.

## 1. Relativistic Coordinate Transformation

An event occurs in a moving frame $S^{\prime}$ at coordinates $\left(x^{\prime}, y^{\prime}, z^{\prime}, c t^{\prime}\right)=(3 m, 1 m, 0,2 m)$. $S^{\prime}$ moves at velocity $v=0.8 c$ in the x-direction with respect to a rest frame $S$. Find the coordinates of the event in $S$.
(A) $(7.66,1,0,7.33)$

Here one need simply apply the Lorentz Transformations:

$$
\begin{array}{r}
x=\gamma\left(x^{\prime}+\beta c t^{\prime}\right)=\frac{5}{3}(3+0.8 * 2)=7.66 \\
y=y^{\prime}=1 \\
z=z^{\prime}=0 \\
c t=\gamma\left(c t^{\prime}+\beta x^{\prime}\right)=\frac{5}{3}(2+0.8 * 3)=7.33 \\
\Rightarrow(x, y, z, c t)=(7.66,1,0,7.33)
\end{array}
$$

## 2. Length Contraction

If a moving observer holds a rod of length $L_{0}=1 \mathrm{~m}$ at an angle of $\theta=45^{\circ}$ from the x -axis, and is travelling at velocity $v=\sqrt{1 / 2} c$, find the length as measured by a stationary observer.
(B) $\sqrt{5 / 8} \mathrm{~m}$

To do this problem one must recall that length conctraction only occurs along the direction of motion. This means that one must break up the bar's length into parallel and perpendicular components. The perpendicular component is unchanged:

$$
\begin{equation*}
L_{\text {perp }}=L_{0, \text { perp }}=L_{0} \sin \theta=1 \mathrm{~m} / \sqrt{2} \tag{1}
\end{equation*}
$$

while the parallel component is contracted:

$$
\begin{equation*}
L_{\|}=L_{0, \|} / \gamma=L_{0} \cos \theta / \gamma=1 m / 2 \sqrt{2} \tag{2}
\end{equation*}
$$

Thus the total length is determined by the two:

$$
\begin{equation*}
L=\sqrt{L_{\text {perp }}^{2}+L_{\|}^{2}}=\sqrt{5 / 8} \mathrm{~m} \tag{3}
\end{equation*}
$$

## 3. Time Dilation

If a moving observer sees a time of $t_{0}=3 s$ elapse on his own watch and a stationary observer measures $t=5 s$ for the same time interval, find the relative velocity of the two observers.
(E)0.8c

Here one must use the formula for time dilation:

$$
\begin{aligned}
t=\gamma t_{0} \Rightarrow \gamma & =t / t_{0} \\
\Rightarrow \frac{1}{\sqrt{1-(v / c)^{2}}} & =5 / 3 \\
\Rightarrow v & =0.8 c
\end{aligned}
$$

## 4. Relativistic Velocity Addition

A particle in rest frame $S$ moves at velocity $\vec{u}=\left(0.8 c, 60^{\circ}\right)$. Find the particles velocity $u^{\prime}$ in frame $S^{\prime}$ which moves at a velocity $\vec{v}=v \hat{x}$ with respect to $S$.
(D) $\left(\mathbf{0 . 7 7 5 c}, 110^{\circ}\right)$

It is first useful to determine the velocity components in frame $S$ :

$$
\begin{array}{r}
u_{x}=0.8 c \cos \left(60^{\circ}\right)=0.4 c \\
u_{y}=0.8 c \sin \left(60^{\circ}\right)=0.4 \sqrt{3} c
\end{array}
$$

One can then find the velocity components in $S^{\prime}$ by plugging these values into the relativistic velocity addition equations:

$$
\begin{array}{r}
u_{x}^{\prime}=\frac{u_{x}-v}{1-\frac{u_{x}}{c} \frac{v}{c}}=\frac{0.4-0.6}{1-0.4 * 0.6}=-0.263 c \\
u_{y}^{\prime}=\frac{u_{y}}{\gamma\left(1-\frac{u_{x}}{c} \frac{v}{c}\right)}=\frac{0.4 \sqrt{3}}{1.25 *(1-0.4 * 0.6)}=0.729 c
\end{array}
$$

from which one can find the new velocity and trajectory:

$$
\begin{equation*}
\left(v^{\prime}, \theta^{\prime}\right)=\left(0.775 c, 110^{\circ}\right) \tag{4}
\end{equation*}
$$

## 5. Relativistic Collision

A particle with rest mass $m_{0}$ travels at a velocity $v \approx c$ with relativistic factor $\gamma_{v}=100$ and collides with a second particle of mass $m_{0}$. If the two particles stick together upon collision, find the factor $\gamma_{u}$ for the velocity $u$ with which the particles travel after the collision.
(C) $\gamma_{\mathbf{u}}=50$

The key to this problem is to use conservation of relativistic momemtum. The momentum of the system before the collision is given by:

$$
\begin{equation*}
p_{i}=\gamma_{v} m_{0} v \approx \gamma_{v} m_{0} c \tag{5}
\end{equation*}
$$

where the approximation $v \approx c$ is used because $\gamma_{v}=100$ is so large. The momentum after the collision is:

$$
\begin{equation*}
p_{f}=\gamma_{u} * 2 m_{0} * u \tag{6}
\end{equation*}
$$

So far, we are not sure if we can approximate $u \approx c$, so we will try to solve for $u$ and then find $\gamma_{u}$. To do this, equate the initial and final momentum (and use the notation $\beta=u / c)$ :

$$
\begin{array}{r}
\gamma_{v} m_{0} c=\frac{2 m_{0} u}{\sqrt{1-(u / c)^{2}}} \\
\gamma_{v}=\frac{2 \beta}{\sqrt{1-\beta^{2}}} \\
\frac{\gamma_{v}^{2}}{4}=\frac{\beta^{2}}{1-\beta^{2}} \\
\beta^{2}=1-1 /\left(\gamma_{v} / 2\right)^{2} \approx 1
\end{array}
$$

Thus we have shown $u \approx c$, so we can now go back to our momentum conservation equation to find $\gamma_{u}$ :

$$
\begin{array}{r}
\gamma_{v} m_{0} c=\frac{2 m_{0} u}{\sqrt{1-(u / c)^{2}}} \approx 2 \gamma_{u} m_{0} c \\
\Rightarrow \gamma_{u} \approx \gamma_{v} / 2=50
\end{array}
$$

